## **Electron Cooling by Vibrational Excitation of Carbon Dioxide**

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## Abstract

We report new results for electron energy transfer rates [1] and electron cooling rates due to vibrational excitation of carbon dioxide (CO<sub>2</sub>). The present model calculations are topical because at altitudes less than about 200 km, the neutral atmospheres of Mars and Venus are dominated by CO<sub>2</sub>. Thus we expect electron cooling by CO<sub>2</sub> to be a significant energy transfer mechanism in their respective atmospheres, with vibrational excitation being the dominant process [2].

Our method was described earlier in Campbell *et al.* [1], although in this application our integral cross section (ICS) database has been meticulously constructed from the best available experimental and theoretical [e.g. 3-5] ICS for vibrational excitation in  $CO_2$ . Electron energy transfer rates from the present work are compared against those from the only previous corresponding study [2], with these results being presented at the meeting.

## References

- L Campbell, M J Brunger, D C Cartwright and P J O Teubner 2004, Planetary and Space Science 52, 815-822.
- [2] M A Morrison and A E Greene 1978, J Geophys Res 83, 1172-1174.
- [3] M J Brunger, S J Buckman and M T Elford 2003, "Photon and Electron Interactions with Atoms, Molecules and Ions", (Springer-Verleg, Berlin), 118-201.
- [4] M Allan 2001, Phys Rev Lett **87**, 033201.
- [5] C W McCurdy, W A Isaacs, H-D Meyer and T N Rescigno 2003, Phys Rev A 67, 042708 (and references therein).